

# Global Positioning System (GPS)



MIEET

1<sup>o</sup> ano



UAlg

UNIVERSIDADE DO ALGARVE



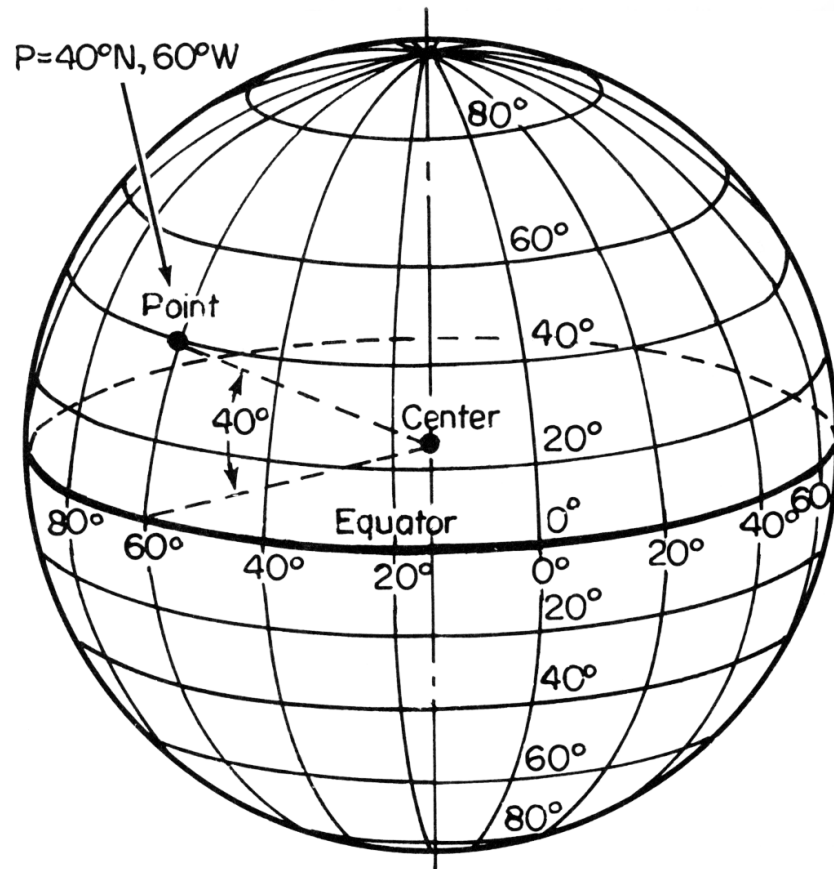
# Why knowing position?



Knowing your position is needed for when you want to go somewhere (and don't get lost)

This is a seemingly simple problem that is (surprisingly) difficult!

# Global coordinates



Position consists of:

- **Longitude** (East-West)  
0° is Greenwich (London)
- **Latitude** (North-South)  
0° is equator, 90° is N/S pole

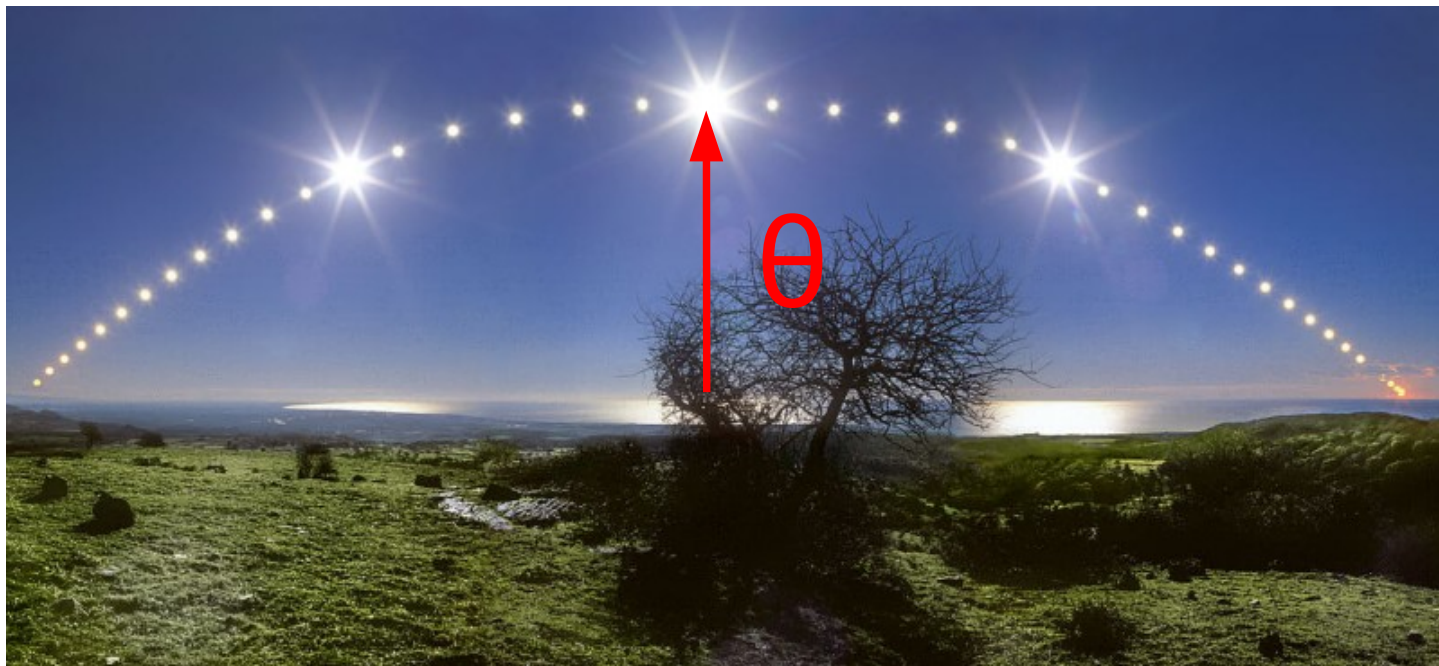
Example: Faro (airport)  
37°00'52"N 007°57'57"W



# Latitude

- Latitude (North-South)

can be determined by height (inclination,  $\theta$ ) of sun in sky (at its highest point!)



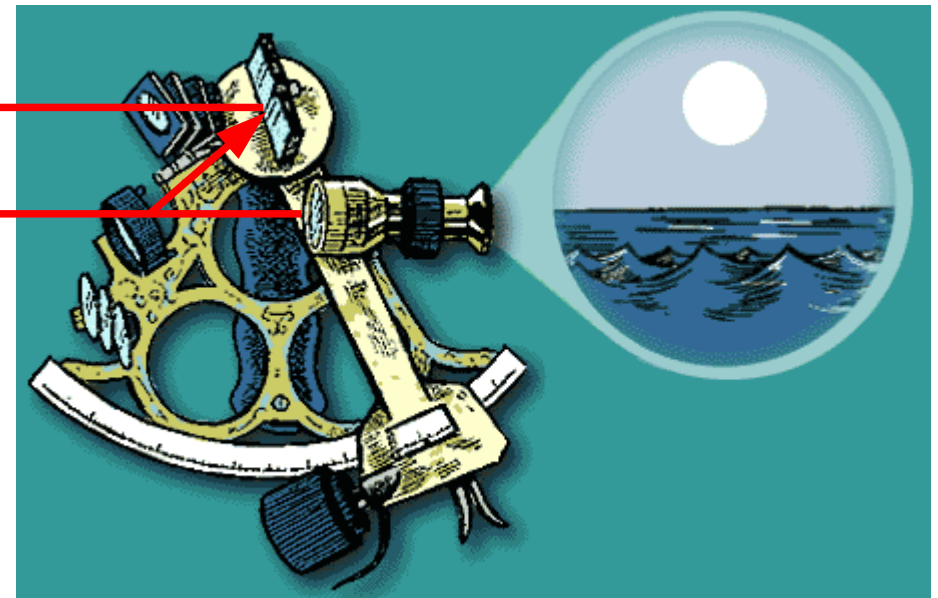
Northern hemisphere (ex. Portugal): left to right →  
Southern hemisphere (ex. Australia): right to left ←



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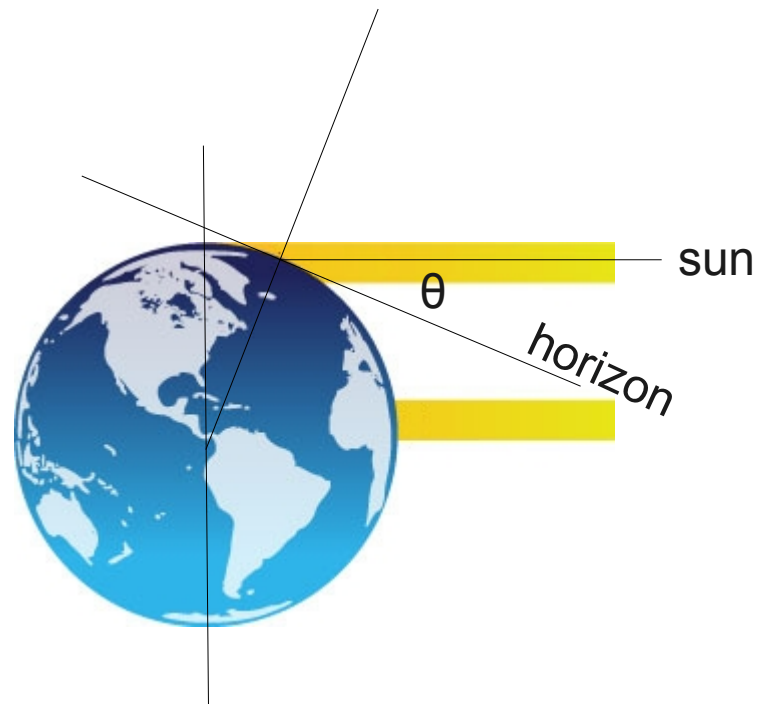


When sun is aligned with horizon, on the scale we can read the inclination

# Latitude

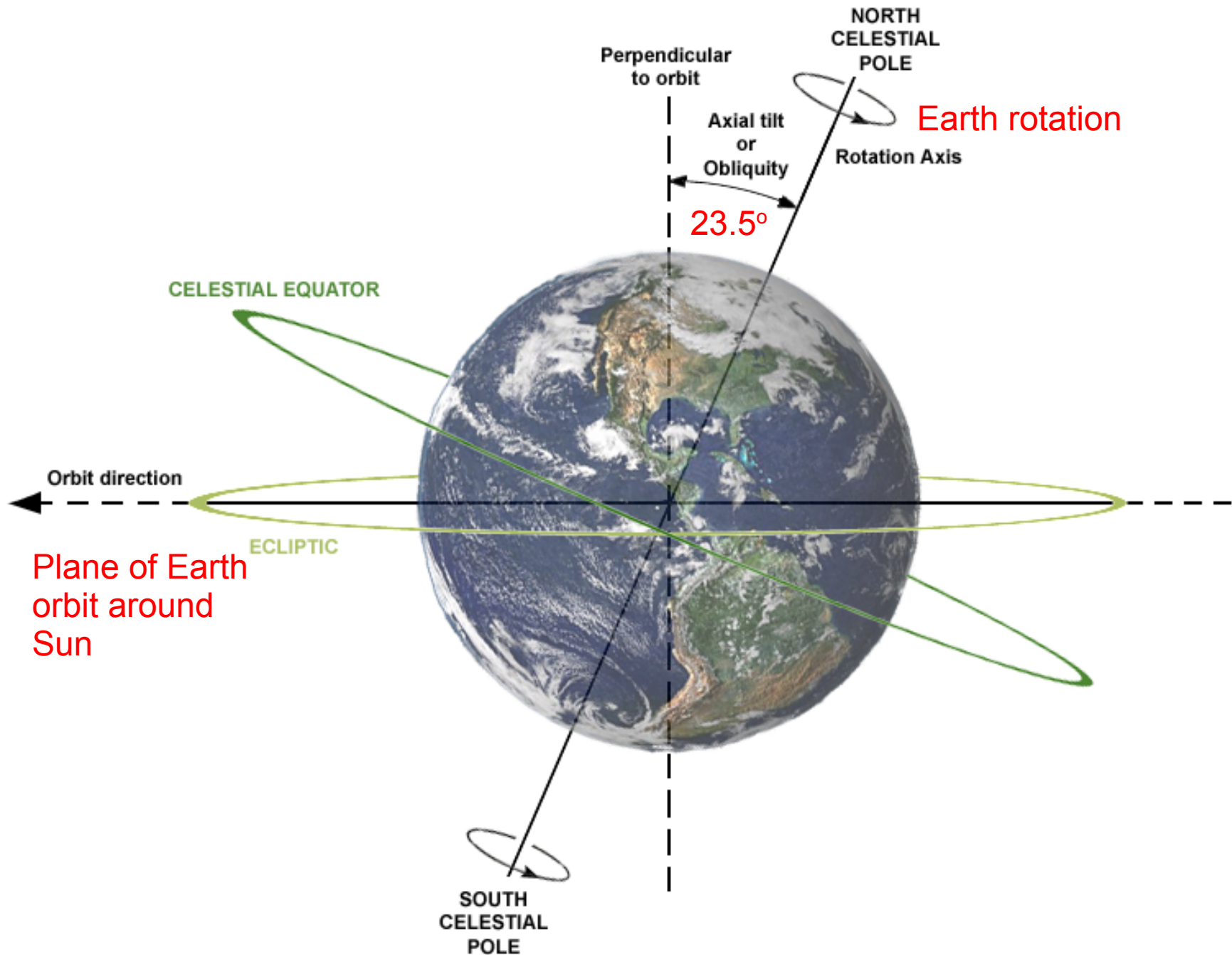
- Latitude (North-South)

can be determined by height (inclination,  $\theta$ ) of sun in sky (at its highest point! at midday)



Simple idea: Latitude =  $90^\circ - \theta$ ?  
More complicated than that!

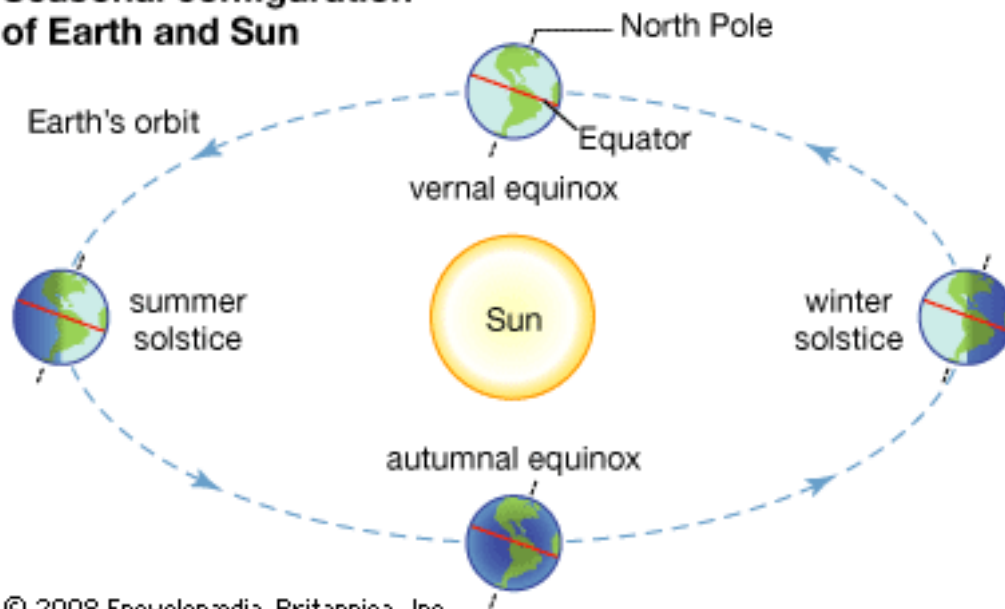
# Latitude; Earth rotation axis



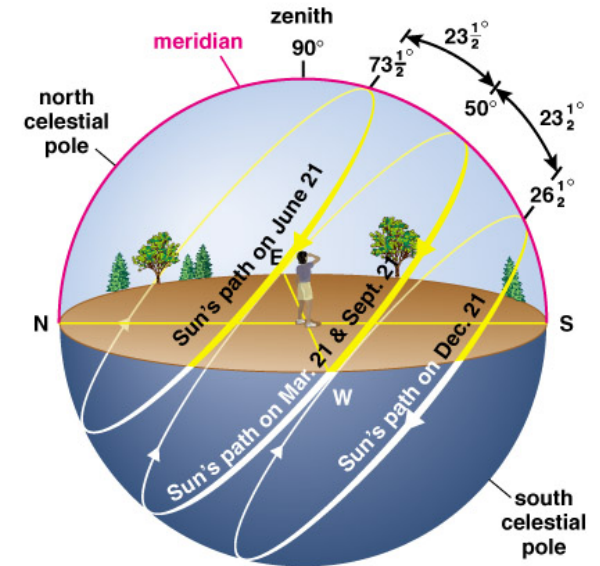
# Latitude

The sun's inclination depends not only on time of day, but also on the day of the year.

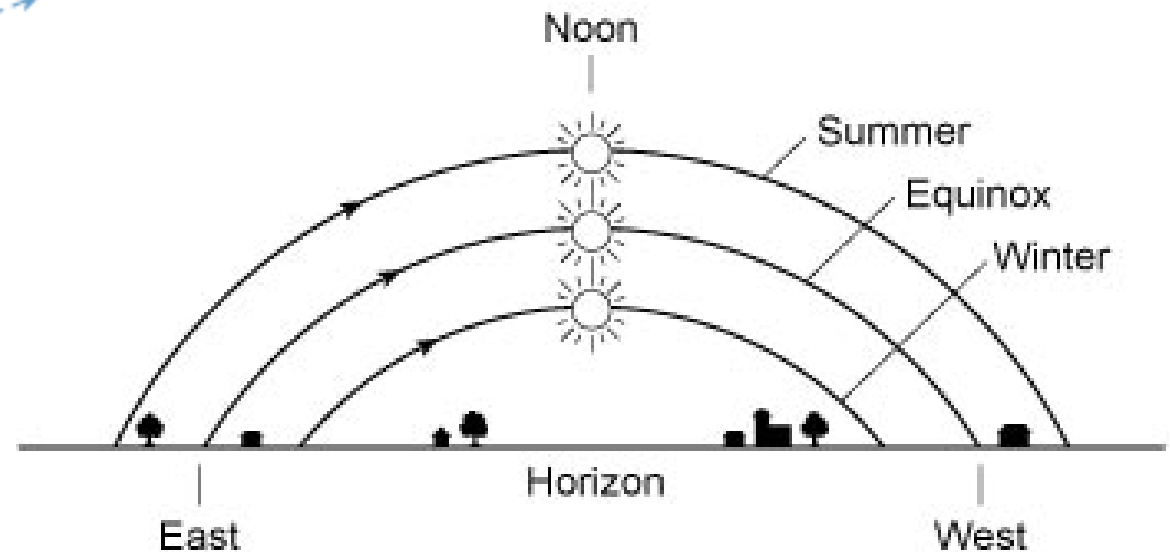
Seasonal configuration of Earth and Sun



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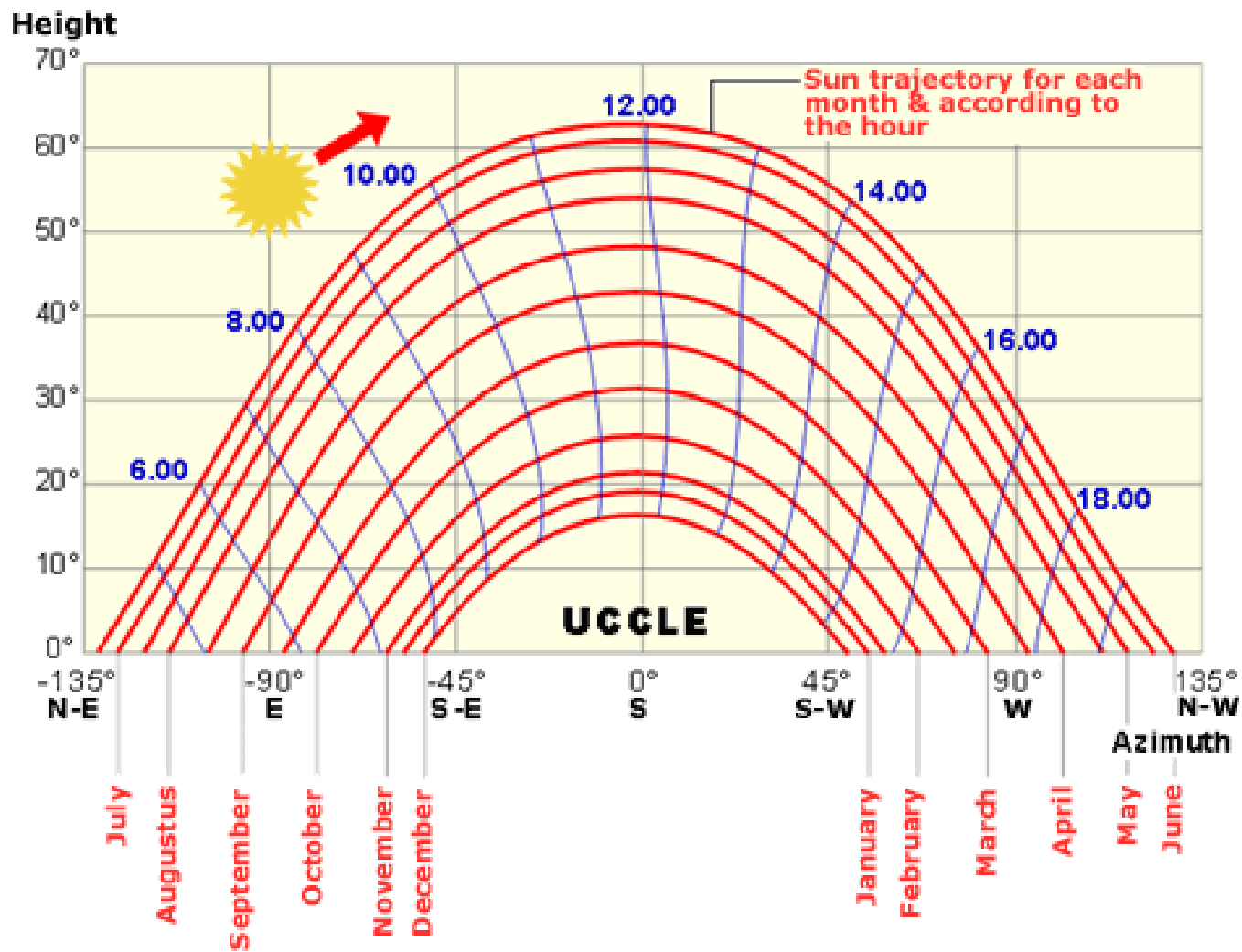


# Latitude

- Latitude (North-South)

Very complicated indeed.

- Sun at highest point **not South**
- Sun **not** always at **same time** of day at highest point

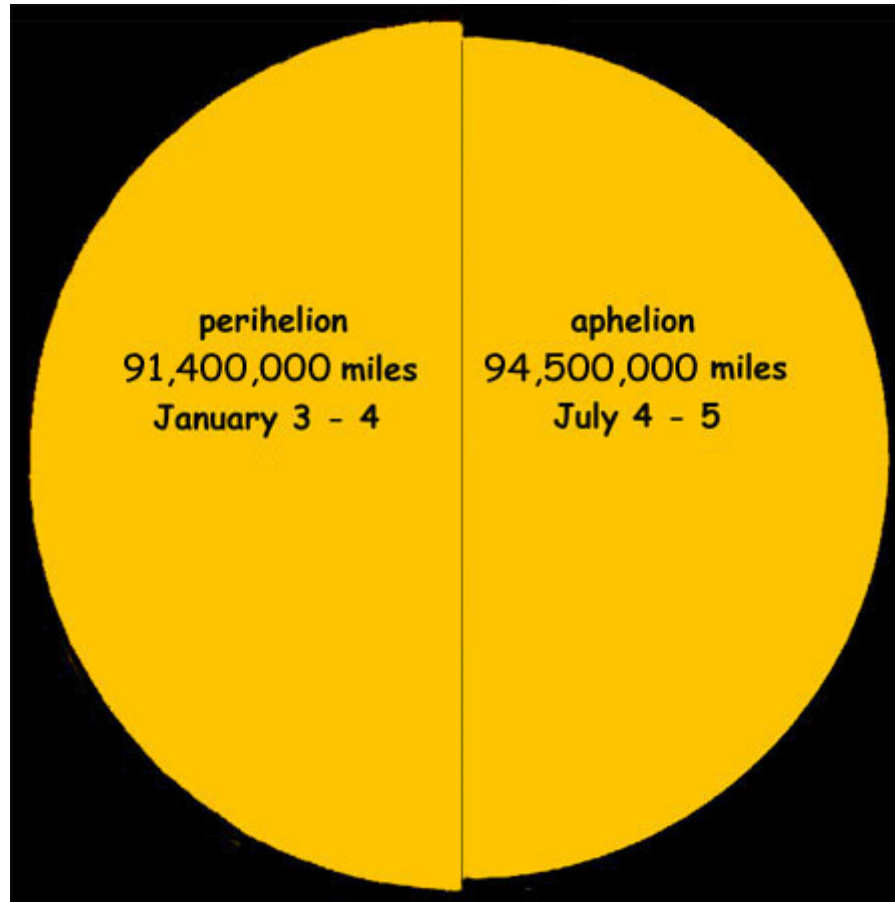


# Latitude



Position of sun at 12:00 along the year  
(Analemma)

# Latitude



The sun is not even the same size along the year.  
Very difficult/unreliable object to use in positioning (and navigation)

# Latitude

- Latitude (North-South)

Better use distant stars. They don't suffer from seasonal effects.

They still change during a day (trajectory).

**Except one!**



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**Except one!**

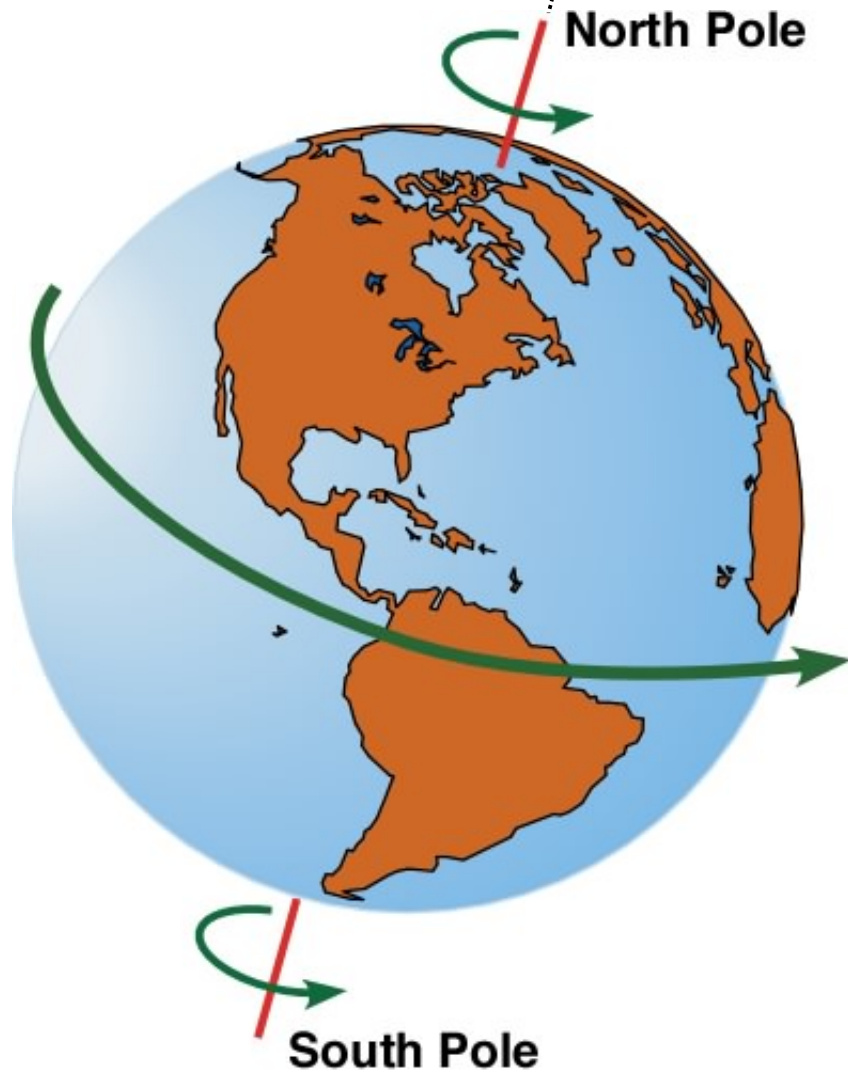
The Polar Star



Long-time exposure at night

# Latitude: Polar Star

\* Polar Star

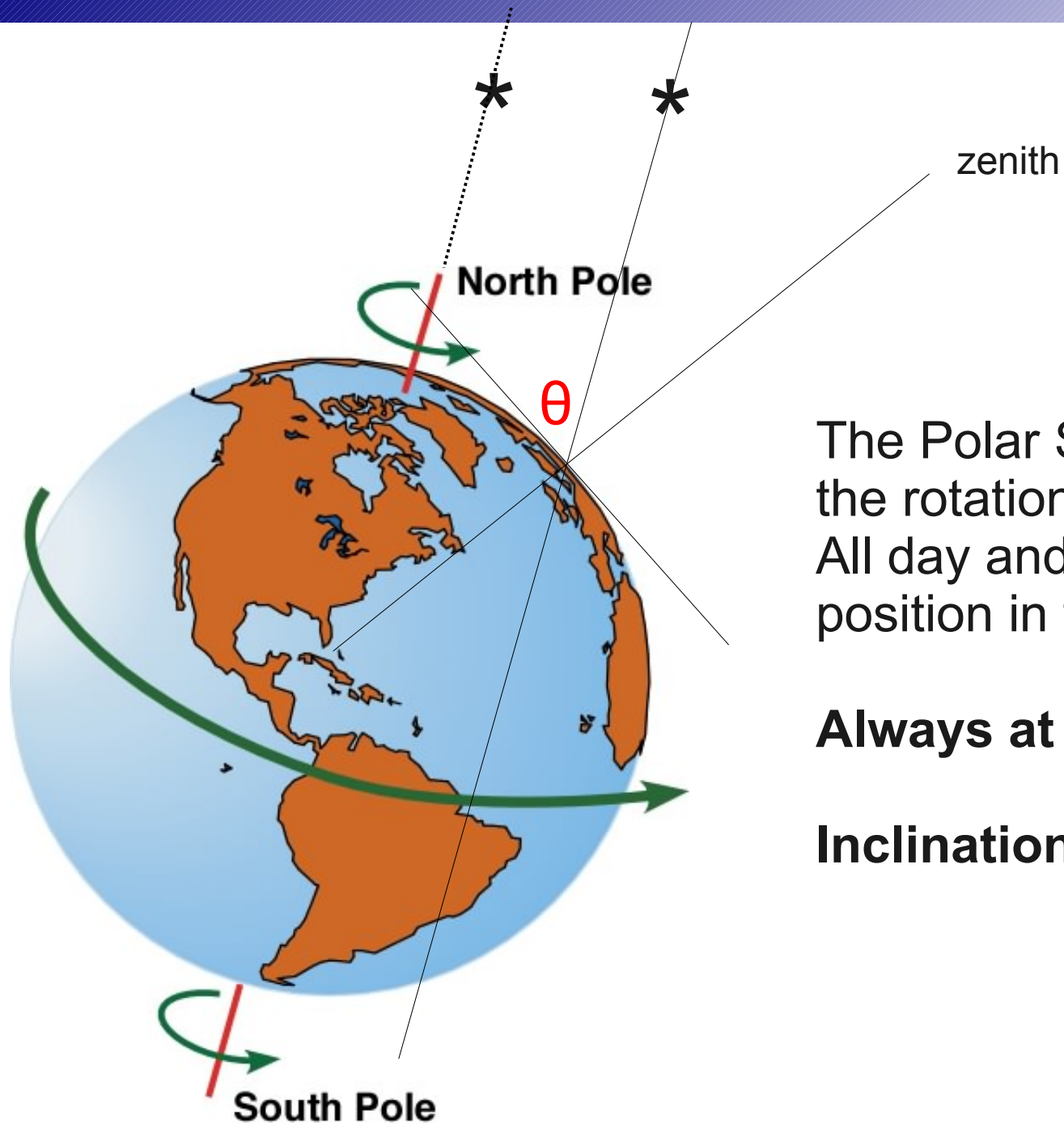


The Polar Star lies on the axis of the rotation of the Earth. All day and all year at the same position in the sky.

**Always at geographic North**

**Inclination equal to latitude**

# Latitude: Polar Star

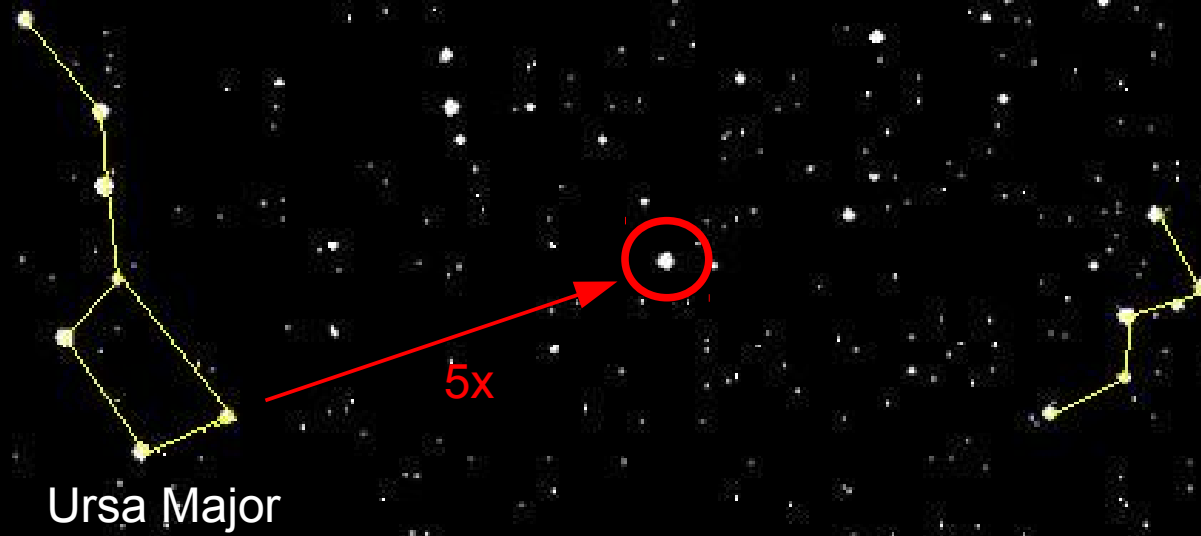


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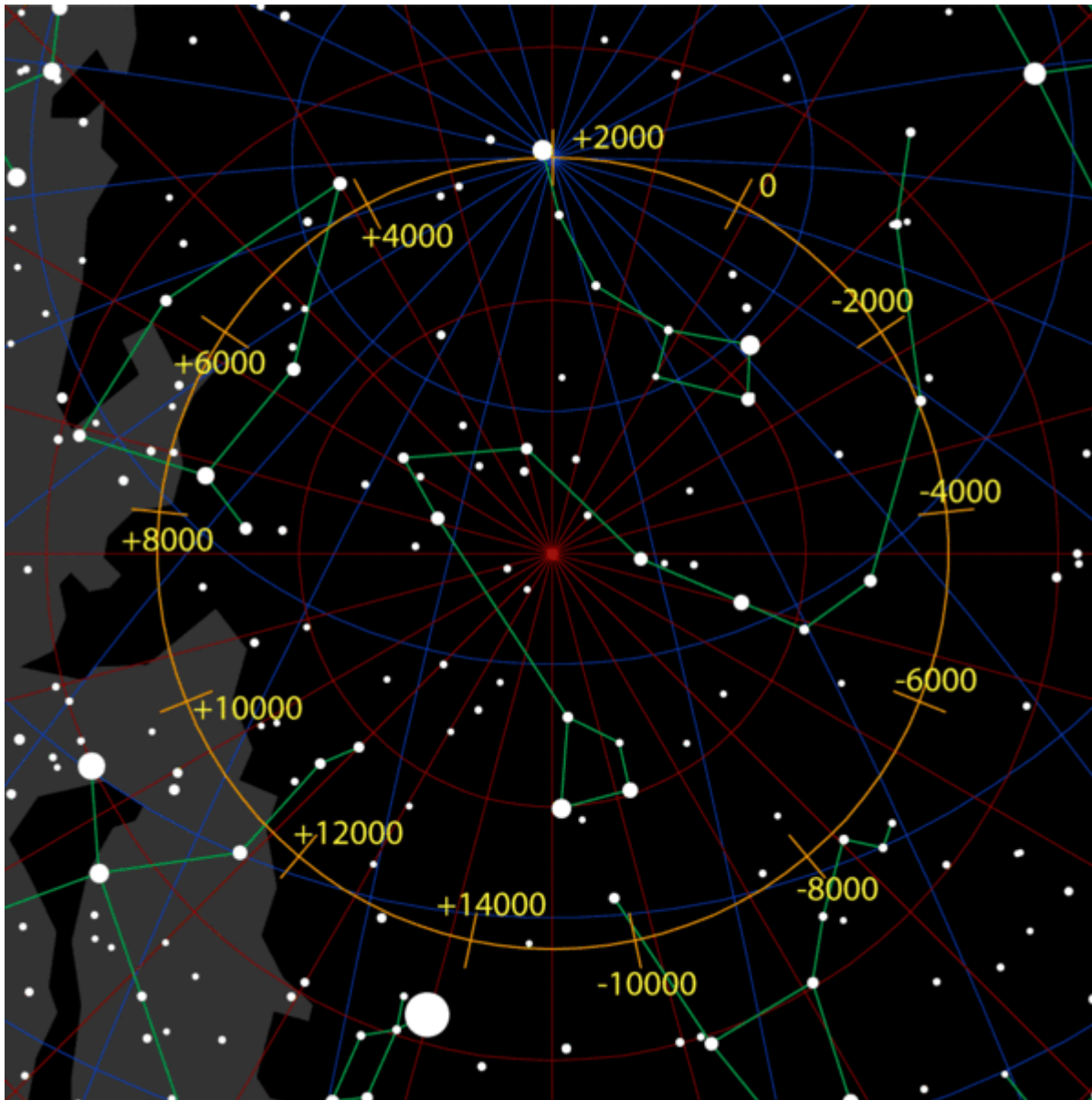
**Inclination equal to latitude**

# Latitude: Polar Star





# Latitude: Polar Star

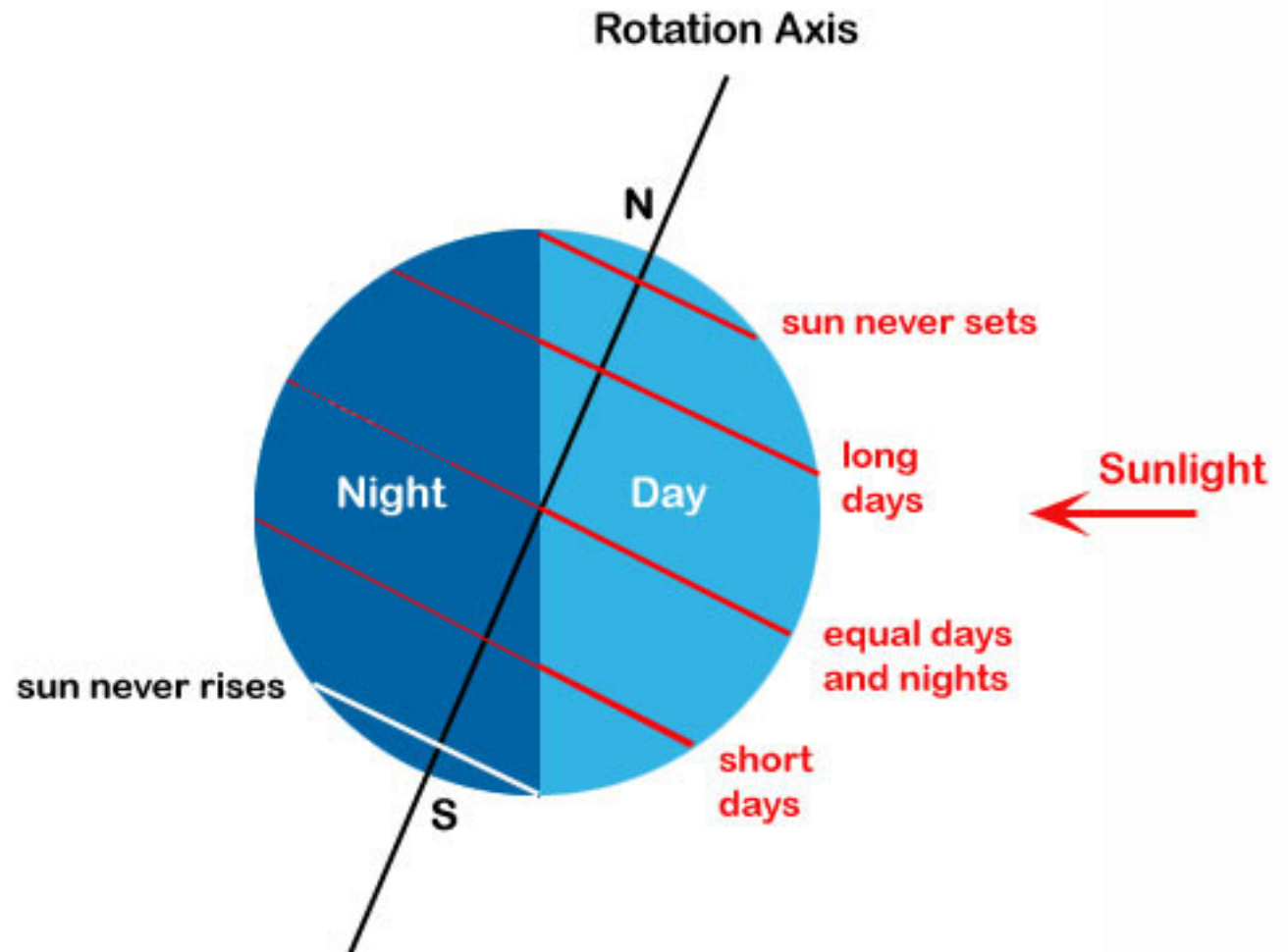


Not even Pole Star  
position is fixed!

Precession of Earth

# Latitude: Polar Star

So far we have just found only the latitude.  
The **easiest** part!



# Longitude

Longitude? Time!

In Greenwich, the sun is at its highest point at exactly 12:00

West: 1 hour later for every  $15^\circ$  ( $360^\circ/24$  h)

East: 1 hour earlier for every  $15^\circ$

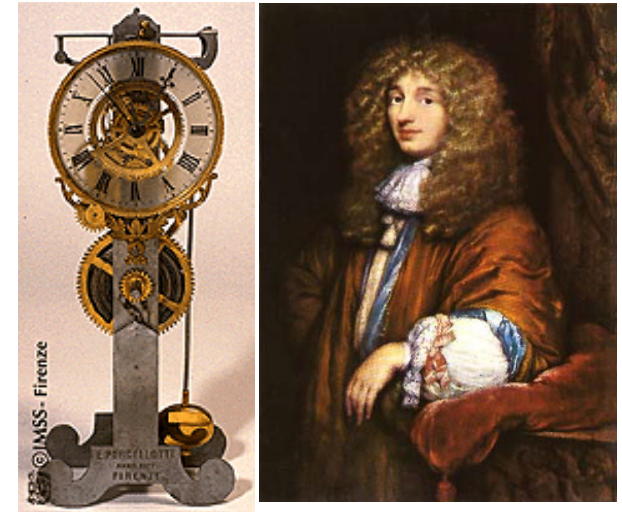
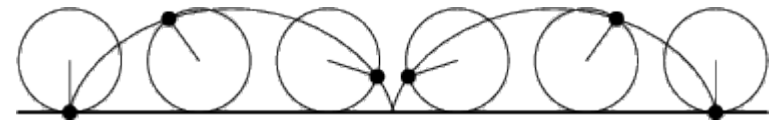
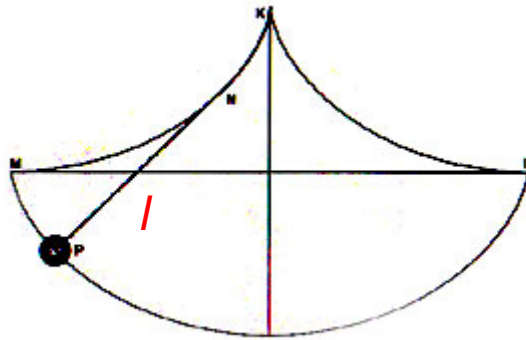
If we know the **time** (GMT) when the sun is at its highest, we know the **longitude**!

The need for precise clocks!

# Longitude: Pendulum clock

Period ( $T=1/f$ ) of a pendulum:

$$T = 2\pi\sqrt{\frac{l}{g}}$$



Period is independent of amplitude of swing if the curve is a cycloid (Christiaan Huygens, 17<sup>th</sup> century)

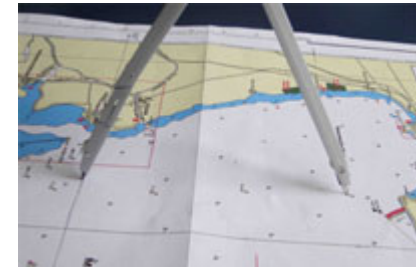


# Comments

With a magnetic compass we cannot measure position, only direction

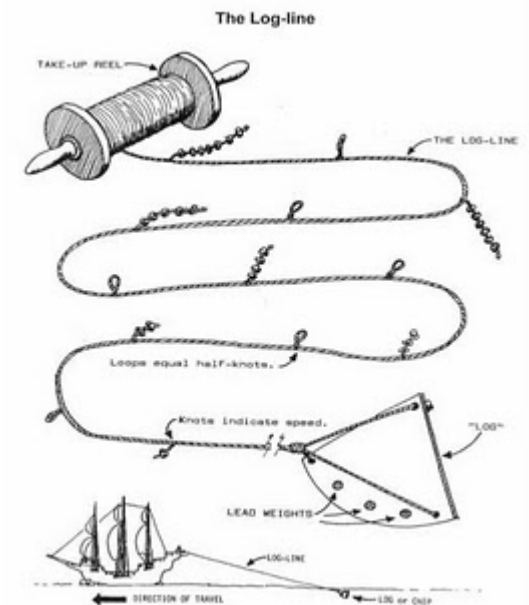


With a mechanical compass we can plot distance on a map (distance is speed x time)



Speed was measured with a rope with knots measure how many equally spaced knots (47 feet, 3 inch) are dragged into the water in 30 seconds

1 knot is equivalent to one nautical mile per hour  
1.852 km/h



# GPS v. 1.0

So far we are still in the preindustrial era, but we know more-or-less our position on this globe.

... and on the Southern Hemisphere? (no North Pole Star!)

... and what if it is cloudy?!!!

GPS.

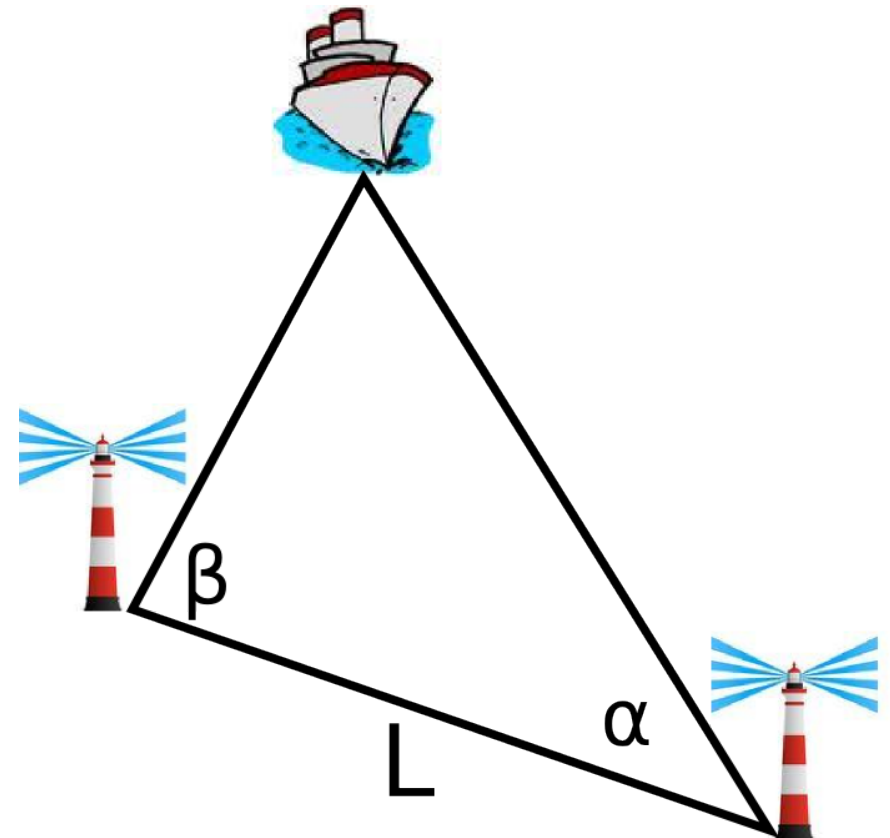
# Angle

If we know the (distance) or **angle** to fixed objects on the globe we can also know our position.

Triangulation is the process of determining the location of a point by measuring angles to it from known points at either end of a fixed baseline.

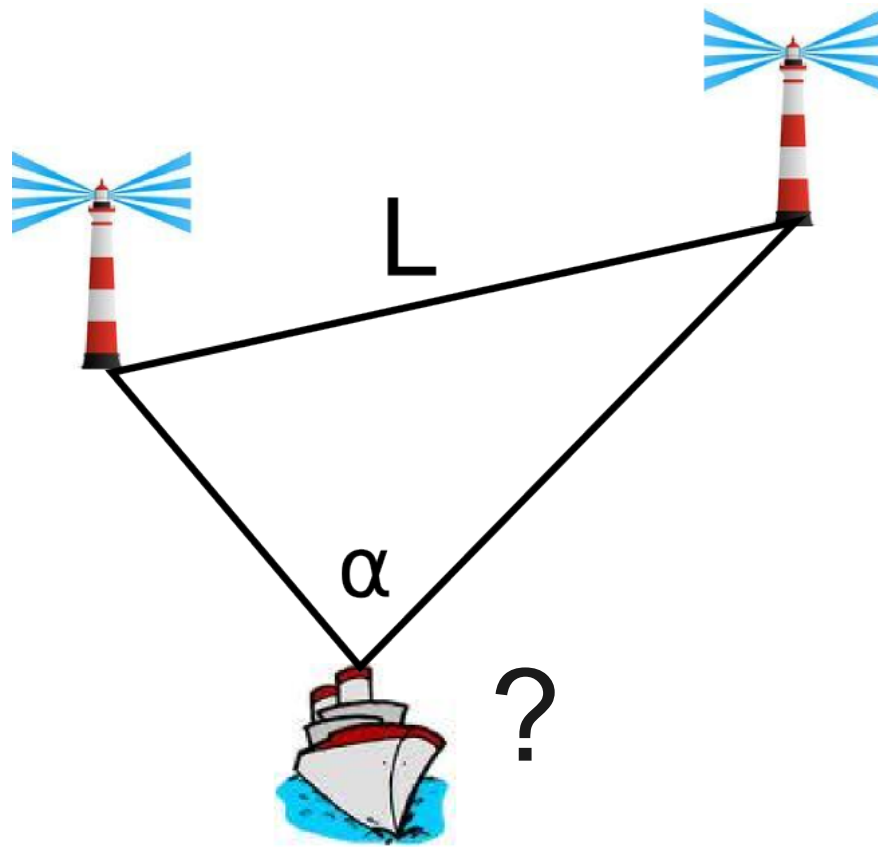
This is based on **trigonometry** (know three things of a triangle and you know everything of the triangle).

For example:  $l$ ,  $\alpha$ ,  $\beta$ . We know the distance to the ship.



# Angle

Note that we know where the ship is, but the ships passengers do not know where they are themselves: Lack of information!

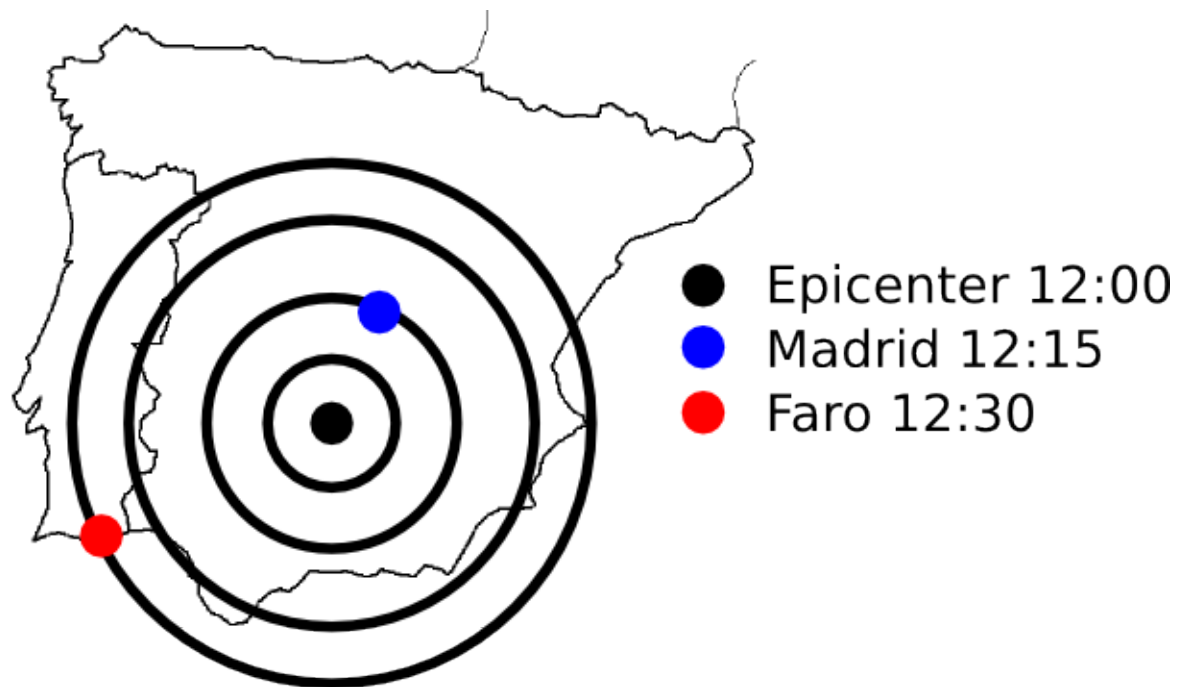




# Distance

If we know the **distance** (or angle) to fixed objects on the globe we can also know our position.

Example: How is the epicenter of an earthquake measured?



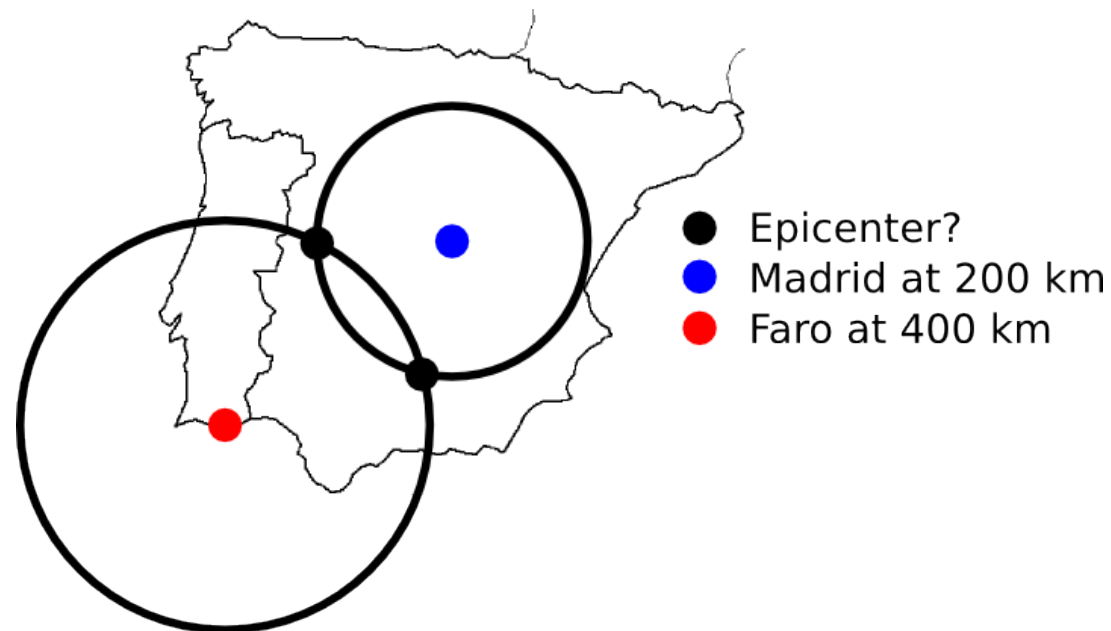
# Distance; Earthquake

Knowing that seismic waves travel at 5 km/s, knowing the time of arrival of the waves tells us the distance.

Madrid: 40 s later = 200 km

Faro: 80 s later = 400 km

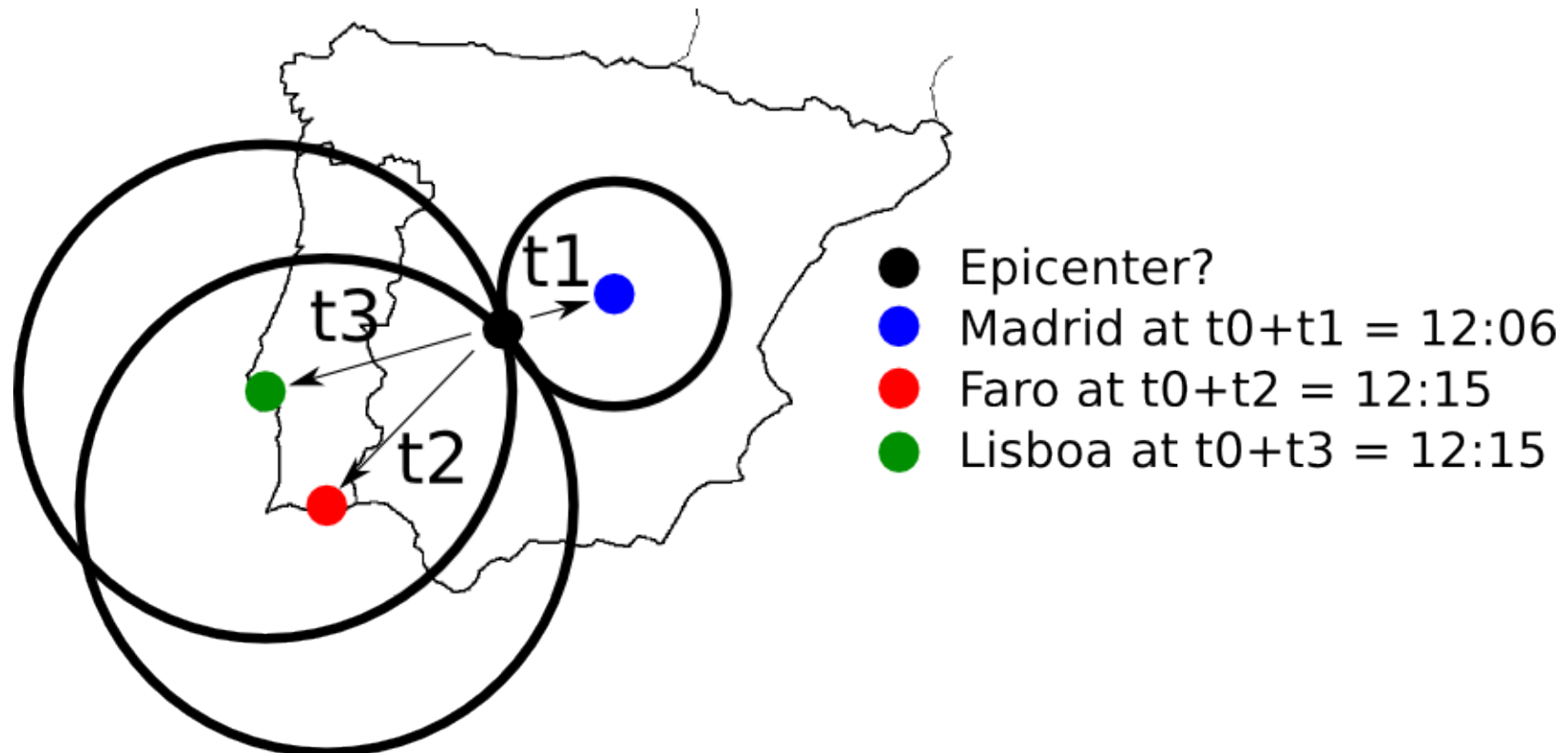
Two solutions:



Apart from that: We need at least 3 (!) stations to determine the epicenter. We don't know **when** the earthquake happened

# Distance; Earthquake

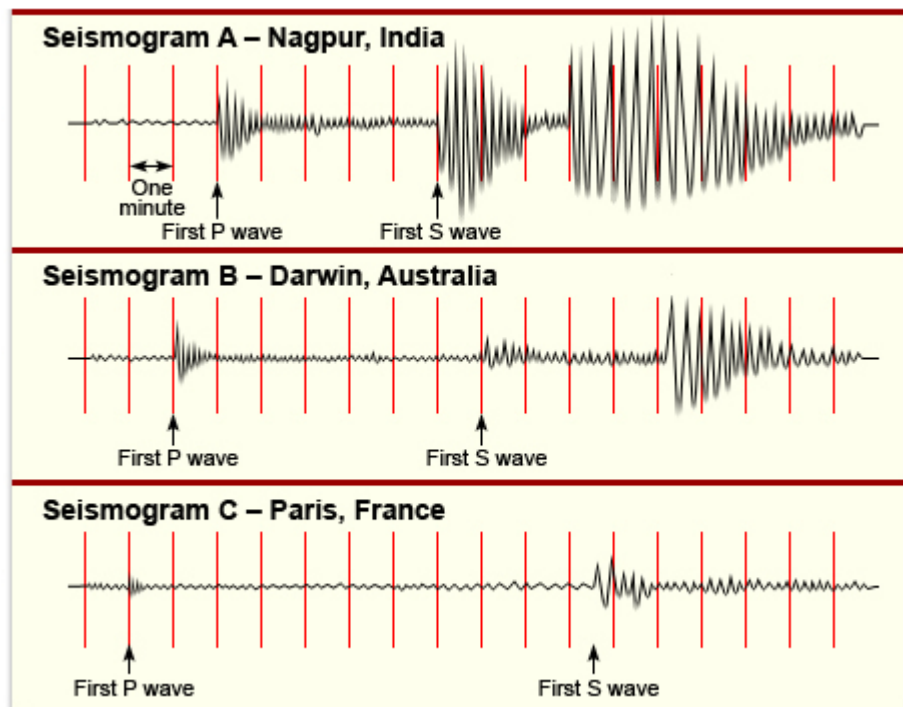
Knowing that seismic waves travel at 5 km/s, knowing the time of arrival of the waves tells us the distance.



We need at least 3 stations to determine the epicenter. Then we also know **when** the earthquake happened

# Distance; Earthquake

In practice difference between arrival of S-waves and P-waves is used (S-waves are transverse, P-waves are longitudinal [like sound waves])



<http://www.digitalgeology.net/>



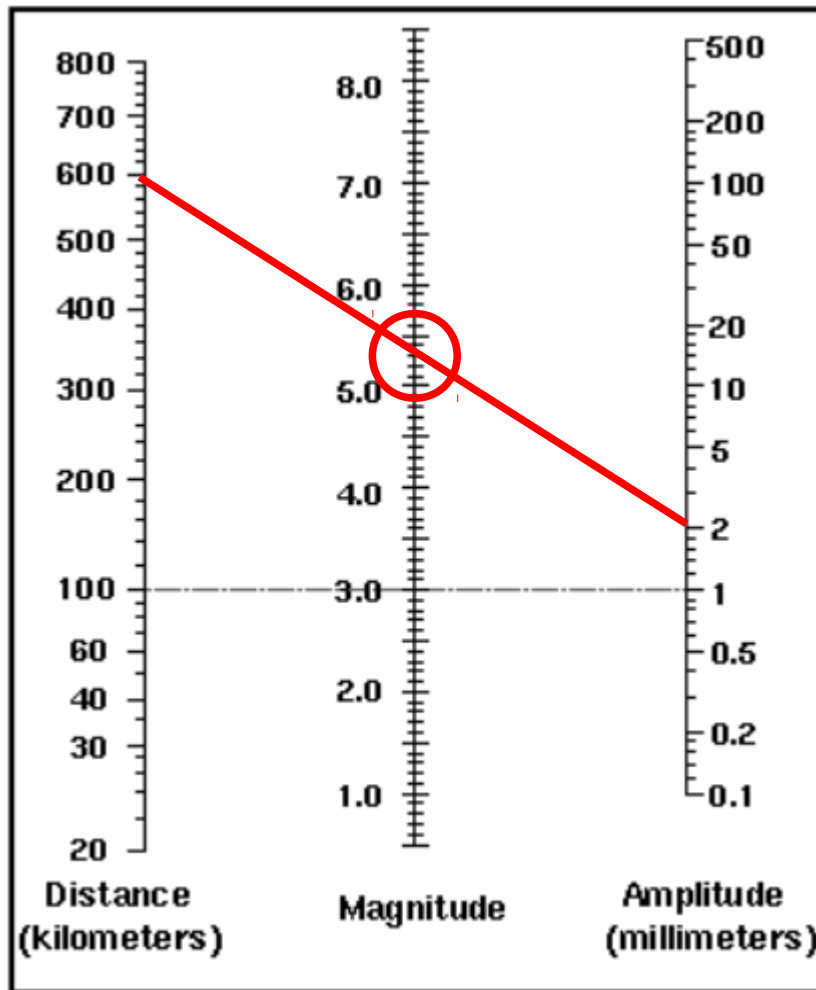
[http://www.tasaclips.com/animations/P\\_S\\_wave\\_velocity.html](http://www.tasaclips.com/animations/P_S_wave_velocity.html)

P-waves: 5 km/s  
S-waves: 3 km/s

The difference in arrival of P-waves and S-waves tells us directly the distance

# Distance; Earthquake

We can also determine the **magnitude** on the Richter scale



Richter nomogram

Example:  
At 600 km, a 5.4 Richter-scale earthquake gives 2 mm amplitude of ground movement

$$A = (1 \mu\text{m}) \times 10^M \times (100 \text{ km}/D)^{1.86}$$

Logarithmic scale:

A scale-4 earthquake is 10 times stronger than a scale-3 earthquake

Power-Law scale:

10 times further away,  $10^{1.86}$  times weaker

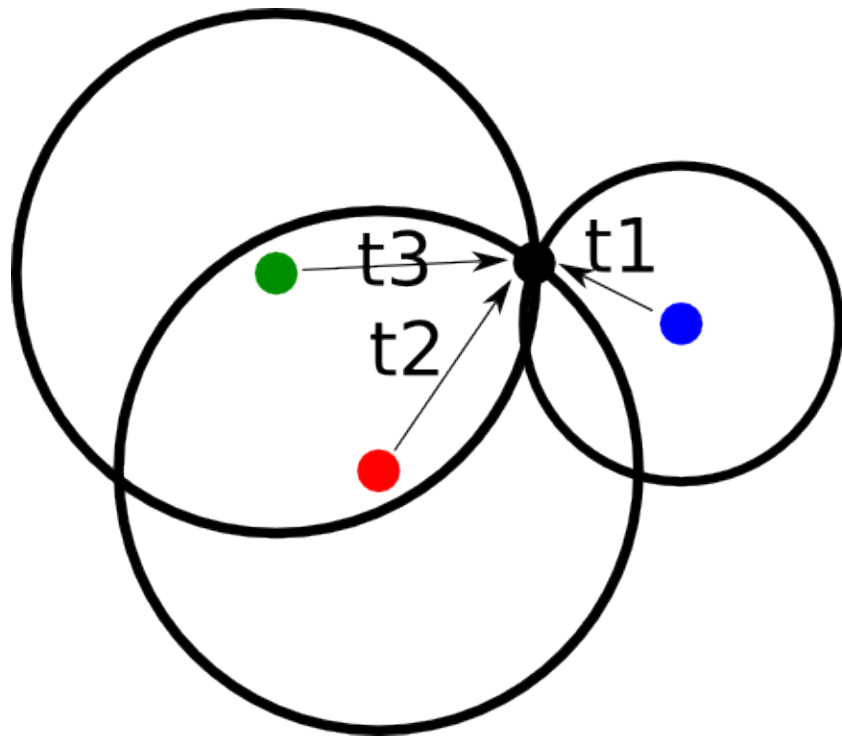


# Own Position; inverted problem

Determining our own position is the inverted problem:

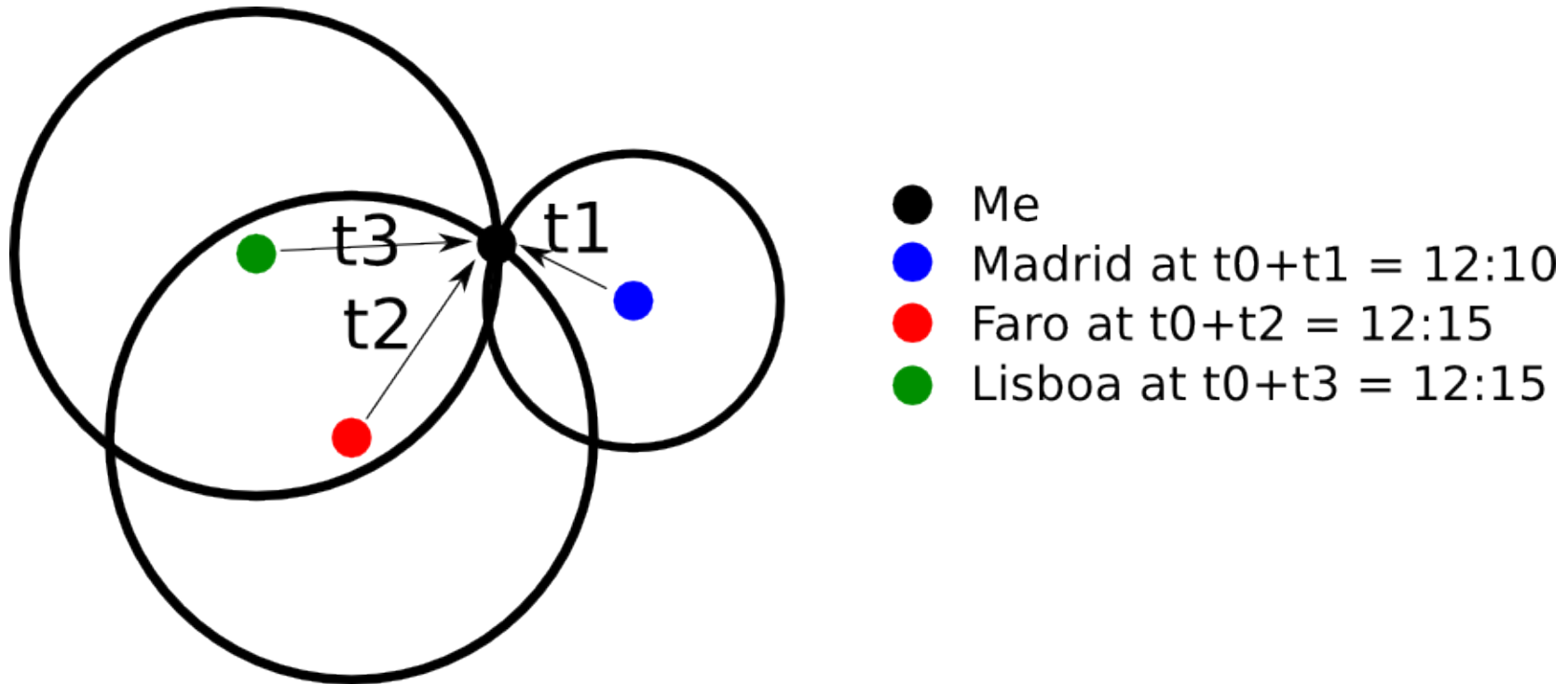
Instead of the epicenter firing in all directions and the different location detecting at different times

All locations firing a signal at same time and I (at epicenter) will receive them at different times



- Me
- Madrid at  $t_0+t_1 = 12:10$
- Faro at  $t_0+t_2 = 12:15$
- Lisboa at  $t_0+t_3 = 12:15$

# Own Position; inverted problem



Positions of stations should be **very** well known

All stations should signal at **exactly** the same time

The arrival time of the signals then tells me the distance to each of the stations and (by trilateration) I can find my own position

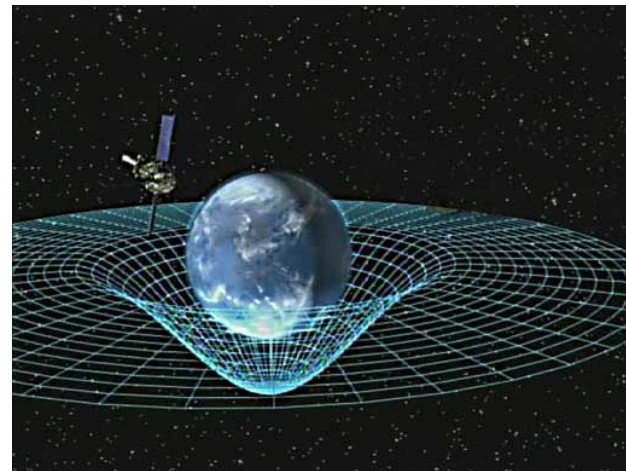
# GPS

GPS works with (24) satellites in orbit around the Earth, which allows for determining **longitude**, **latitude** and **height**

Electromagnetic waves traveling with speed of light,  $c = 3 \times 10^8$  m/s

We want position within, say, 10 m.

Timing should be accurate within  $(10 \text{ m}) / (3 \times 10^8 \text{ m/s}) = 33 \text{ ns}$ .



# GPS, satellites with atomic clocks

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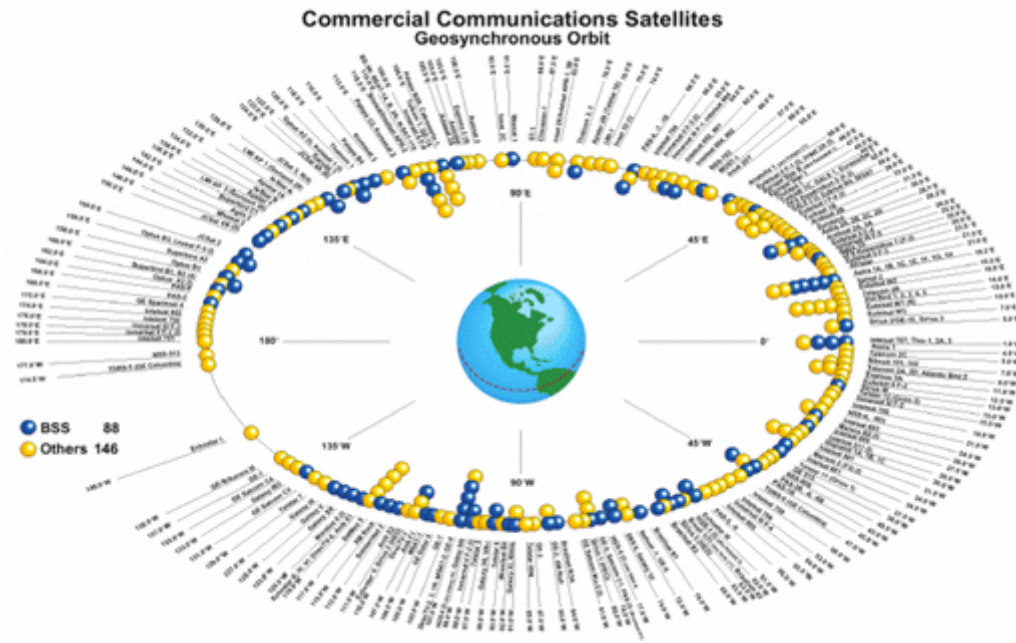
**Atomic clocks** on all satellites. (No atomic clock needed on GPS receiver!), calibrated constantly by a ground station

Relativistic (Einstein) effects important. (In fact, **GPS is the result of a scientific study on relativistic effects**)

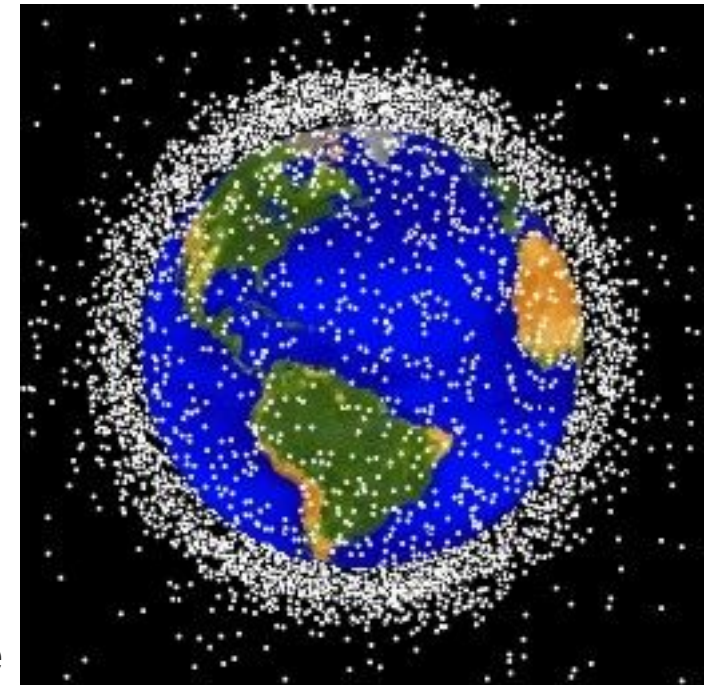
**4 satellites** are enough, but with more (there are 24 today) we get better accuracy

# Satellites in space. Where is Greenpeace?!!

Figure 3: Geostationary Satellites by Orbital Location



Geostationary (communication) satellites



All satellites in Earth space



# One way

In GPS, only the satellites send signals. The receiver not.  
**You do not reveal your location to anybody!**

The position of mobile phones can also be determined by looking at the time delays to/from the cellular towers and the signal strength.

**They know where you are!**



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Whatever you do, **don't**  
bring your mobile phone!!